

More on Boltzmann eq

→ Summary of Basic Boltzmann eq.

- Basic Eqn.

$$C(F) = \int d\underline{p}_2 \frac{\partial V_{12}}{\partial r_1} \cdot \frac{\partial}{\partial p_1} [F(1,t) F(2,t)]$$

$$\text{Collision operator} = \int d\underline{p}_1 \int d\underline{p}' \int d\underline{p}_1' w(\underline{p}', \underline{p}_1'; \underline{p}_1, \underline{p}) (f(\underline{p}') f(\underline{p}_1') - f(\underline{p}_1) f(\underline{p}))$$

with

$$w = w(\underline{p}', \underline{p}_1'; \underline{p}_1, \underline{p}) = w^T \quad (\text{time reversible})$$

as transition probability

⇒ R.E.

$$\frac{\partial f}{\partial t} + \underline{v} \cdot \underline{\nabla} f = C(F)$$

$\left\{ \begin{array}{l} \text{based on} \\ \text{- detailed balance} \\ w^T = w \\ \text{- molecular chaos} \end{array} \right.$

$$\text{Now, } S = - \int d\underline{x} \int d\underline{p} f \ln f \rightarrow \text{entropy.}$$

$$\text{and : } \frac{dS}{dt} = \frac{1}{2} \int d^4 p \int d^3 x w f f_i [x \ln x - x + 1]$$

$$x = f' f_i / f f_i$$

$$\frac{dS}{dt} \geq 0 \rightarrow H \text{ thm.}$$

Max entropy state, $x=1 \Rightarrow dS/dt = 0$.

⇒ Some observations:

①

→ never actually used concept of equilibrium or equilibrium distribution function in building Boltzmann eq, though did observe $C(f_{eq}) = 0$.

→ might ask: "If no a-priori knowledge of equilibrium distribution, could one derive it?"

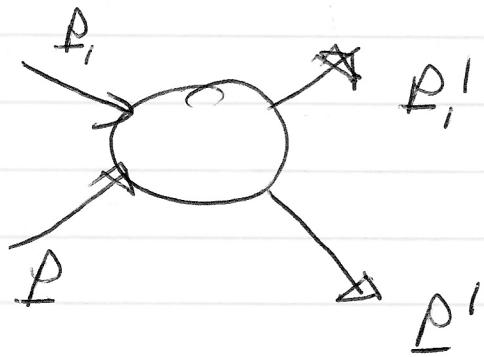
Now:

$$\frac{dS}{dt} = 0 \quad \text{for } x=1 \quad \left\{ \begin{array}{l} \text{Point of} \\ \text{maximal} \\ \text{entropy} \end{array} \right.$$

$$x=1 \Rightarrow f' f'' = f f'_i$$

$$\ln f + \ln f'_i = \ln f + \ln f'$$

as labels in collision are arbitrary,
i.e.



have, at equilibrium:

$$\ln f + \ln f_i = \text{const}$$

\Rightarrow {Sum of logs conserved
in collision.}

{for equilibrium
 $(ds/dt = 0)$ }

Now, what is constant in collision:

{ - energy (kinetic for pt. particle)
- momentum
- number }

\Rightarrow $\ln f$ can be expressed as a linear combination (with constant coeffc) of quantities conserved in a collision,

so

$$\ln f = a + \frac{b}{\text{momentum}} + c \frac{p^2}{2m} + d \frac{\text{KE}}$$

Note: - CGO for normalizability of f

- conservation requires form of $\ln f$

- angular momenta not independent,
as event occurring at 1 position.
collision

$$\Rightarrow f = C' \exp \left[-\frac{p^2}{2mT} - \frac{p \cdot V}{T} \right]$$

\Rightarrow recovers Maxwell-Boltzmann distribution

$\Rightarrow \Lambda, T, V$ all can be functions of x
for $\text{lmp} \ll L$

\Rightarrow have derived form of eqbm distribution
from condition $dS/dt = 0$

⑥

\Rightarrow Boltzmann H-Thm exploited concept of entropy. Where from?

Fundamentally, Entropy $\leftrightarrow \ln (\text{Phase Volume})$

② $S = \ln A\Gamma$ \rightarrow Fundamental definition

$$= \ln \frac{\Delta p \Delta q}{(2\pi\hbar)^5}$$

\hookrightarrow assumed
dimensions.

Now $f(\bar{E}) A\Gamma = 1$

of weight factor \Rightarrow assumes ~~tightly localized f~~
mean energy \Rightarrow ~~tightly localized f~~

and linearity of log:

$$\ln F(E) = \alpha + \beta E \quad (\text{uses struct. of Esbrem})$$

so $\ln f(\bar{E}) = \alpha + \beta \bar{E}$ \Rightarrow can obtain from Liouville Eqn with sub-systems.

\Rightarrow

$$\ln f(\bar{E}) = \alpha + \beta \bar{E} = \langle \ln f(E) \rangle$$

thus

$$S = \ln A\Gamma = -\ln F(\bar{E}) = \ln \langle f(E) \rangle$$

\Rightarrow

$$S = -\langle \ln f(E) \rangle = -\int f \ln f d\Gamma$$

\Rightarrow recovers entropy used in H-Thm.

(c) How Reconcile?

- reversible laws of Hamiltonian mechanics, which govern gas
- $dS/dt \geq 0$.

related:

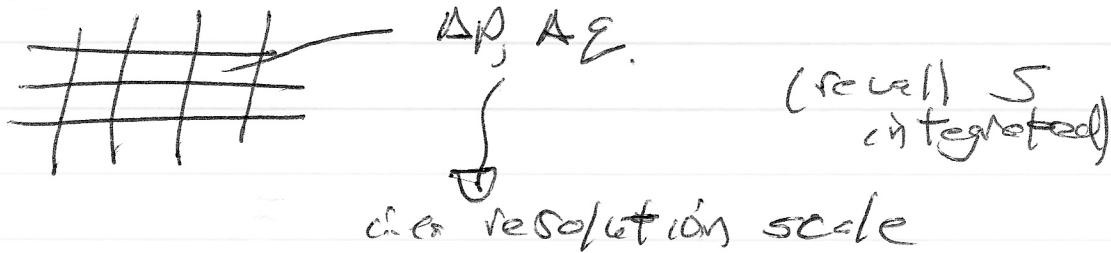
- whatever happened to Poincaré recurrence?

Key Point: Boltzmann introduced:

\rightarrow statistical description
i.e. $F(x, p, t) \rightarrow$ probability

\rightarrow coarse graining (recall Lyapunov exponents!)

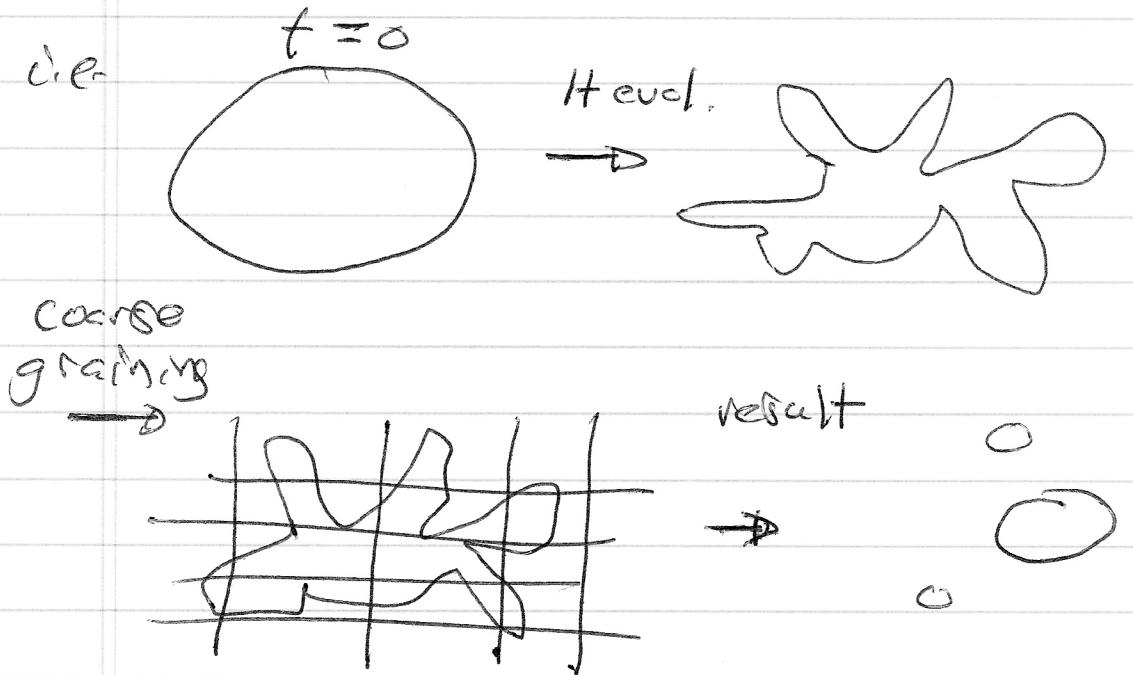
i.e. partition



Z

Why significant?

→ partition / coarse graining kills small details in phase volume evolution



Zapped
(i.e. phase space
reconnection) by
coarse graining.



prediction of very low probability events is
impossible, inaccurate.

and

↯ recurrence is very low probability.

(d) Why is statistical description valid?

→ chaos

(even for $N=2$,
not only $N=N_A$)

$\leq \tau_{\text{relax}} \gg \tau_{\text{typ}}$

mixing time

i.e. calculated $\rho_M \rightarrow$ it works.

→ what are the key assumptions:

→ reversible, conservative collisions

→ $f(l_1, z_1) = f(l_1) f(z_1)$
(Molecular chaos)

[chaos \rightarrow correlation
mixed]

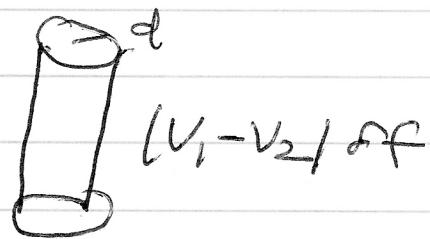
Per Boltzmann: Stosszahlansatz

i.e.

total # (V_1, V) collisions taking place
in δt

= volume of (V_1, V) collision cylinder
(*) # particles with V_1 , per volume.

L
Collision cylinder:



$$\# \text{ in collision cylinder} = V_{\text{rel}} T F(v_i) d^3 v_i$$

No. B.:

- dilute \rightarrow non-overlapping cylinders

- collisions \approx ('point events')

$$d < \bar{r} < l_{\text{mfp}} \quad \text{ordering!}$$

- $f(l_1, l_2)$ factorization

all buried here.

Also

$$\rightarrow w d^3 V' d^3 V' = V_{\text{rel}} dT$$

relates transition to familiar items
like cross-section.

$$\rightarrow \text{for } l_{\text{mfp}}, \quad l_{\text{mfp}} = 1/n \sigma$$

$\tau_{\text{Lmp}} = \text{volume of}$
 $\text{collision cyl., for 1 collision.}$

∴ $N \tau_{\text{Lmp}} = \# \text{coll} = 1$  $\tau_{\text{Lmp}} / N = 1$

$$\boxed{\ell_{\text{mp}} = 1/N}$$

$\nu_{\text{coll}} = v_{\text{th}} / \ell_{\text{mp}}$ \rightarrow defines collision frequency.

a.b. crudely: $\tau \sim d^2$.

Fluid Equations

- what to do with B.E. \Rightarrow

$\begin{cases} \text{fluid eqns.} \\ \text{transport coefficients} \rightarrow \text{flux-force proportionality} \end{cases}$

- replace B.E. by set of equations which evolve thermodynamic parameters.

\Rightarrow hydrodynamic equations for $n(x, t)$, $\bar{v}(x, t)$, $p(x, t)$ etc.

- local, Eulerian description (lab frame) \rightarrow held together by collisions, v.f. dispersion
- describes 'blobs' of gas, held together by collisions, $L > \text{lmfp} \rightarrow$ order of $\omega < r_{\text{coll}} \rightarrow$ relevance
- parametrizes dynamics in terms structure of distribution

$$f = \frac{n(x)}{(2\pi)^{3/2} v_{th}(x)^3} \exp \left[- \frac{(v - \bar{v}(x, t))^2}{v_{th}^2(x, t)} \right]$$

- works for slight deviation from equilibrium

$$\text{i.e. } f = f_0 + \delta f$$

~~loses~~
Maxwellian

~~f~~
ideal
equations

$$\hookrightarrow \delta f = -\frac{\underline{v} \cdot \nabla f_0}{v}$$

~~f~~
viscous
dissipative
eqns.

sacrifices info re: higher moments (useful energy truncation).

now, recall:

$$\frac{\partial f}{\partial t} + \underline{v} \cdot \nabla f = C(f)$$

$$\text{demand: } \int d^3\underline{v} C(f) = 0 \quad (\# \text{ conservation})$$

$$\int d^3\underline{v} m\underline{v} C(f) = 0 \quad (\text{momentum conservation})$$

$$\int d^3\underline{v} E C(f) = 0 \quad (\text{energy conservation})$$

then define:

$$n = \int d^3v f \rightarrow \text{density}$$

$$\underline{v} = \frac{1}{n} \int d^3v \underline{v} f \rightarrow \text{velocity / momentum}$$

$$\bar{E} = \frac{1}{n} \int d^3v \, \epsilon f \rightarrow \text{energy density}$$

and $\frac{\partial}{\partial t} f + \frac{\partial}{\partial x_i} (v_i f) = c(f)$

taking moments:

$$\frac{\partial n}{\partial t} + \nabla \cdot (n v) = 0$$

+ +
mass flux # cons.

$$\frac{\partial}{\partial t} (m v_\alpha) + \frac{\partial}{\partial x_\beta} \Pi_{\alpha\beta} = 0 \quad (\text{momentum balance})$$

+ +
momentum flux mom. cons.

$$\Pi_{\alpha\beta} = \int d^3v \, m v_\alpha v_\beta f$$

$$\frac{\partial(\bar{E})}{\partial t} + \nabla \cdot \bar{Z} = 0 \quad (\text{energy balance})$$

+ +
energy flux energy cons.

$$\bar{Z} = \int d^3v \, v \epsilon f$$

note form:

$$\frac{\partial}{\partial t} (\text{stuff}) + \underline{\partial} \cdot (\text{Flux of stuff}) = 0$$

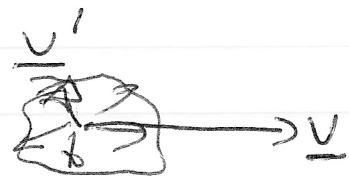
\rightarrow essence is calculation of fluxes.

Further simplify by:

$$\underline{v} = \underline{\bar{v}}(x, t) + \underline{v}'$$

↓
 mean/bulk
 flow
 (~ linked to
 body forces)

↓
 thermal
 fluctuation
 about mean
 ($\sim \sqrt{kTm}$)



realistically!
 $|\underline{v}| \ll |\underline{v}'|$

{ but \underline{v} & \underline{v}' cancel}

$$\Pi_{\alpha, \beta} = \int d^3v m (\underline{v}_\alpha(x, t) + \underline{v}'_\alpha) (\underline{v}_\beta(x, t) + \underline{v}'_\beta) f$$

$$f = f_0 + \cancel{f}$$

loc. Maxwellian, as above.

so

$$\Pi_{\alpha, \beta} = mn (\underline{v}_\alpha(x, t) \underline{v}_\beta(x, t) + \langle \underline{v}'_\alpha \underline{v}'_\beta \rangle)$$

now, $f = \frac{n(x)}{(2\pi)^{3/2} k_B^3(x)} \exp \left[- \frac{(\underline{v} - \underline{\nabla}(x,t))^2}{k_B^2(x)} \right]$

$$\langle \underline{v}'_A \underline{v}'_B \rangle = \frac{1}{3} \underline{v}^2 \delta_{AB} \quad (\text{isotropic } f_0 \text{ off } \underline{v})$$

$$\langle \underline{v}^2 \rangle = 3T/m$$

(collisions
isotropic
pressure)

so, can define:

$$\underline{\underline{P}} = mn \langle \underline{v}'_A \underline{v}'_B \rangle$$

$$= \frac{1}{3} mn \langle \underline{v}^2 \rangle \delta_{AB}$$

$$\underline{\underline{P}} = nT$$

so

$$\frac{\partial}{\partial t} (n\underline{v}) + \underline{\nabla} \cdot (n\underline{v}\underline{v} + \underline{\underline{P}}) = 0$$

but: $\frac{\partial n}{\partial t} + \underline{\nabla} \cdot (n\underline{v}) = 0$

subtracting \Rightarrow

$$n \left(\frac{\partial \underline{v}}{\partial t} + \underline{v} \cdot \underline{\nabla} \underline{v} \right) = -\underline{\nabla} P$$

"Euler Eqn."
(+ visc \rightarrow Navier-Stokes).

$\nabla \cdot \underline{V} = 0$ defines P (incompressible fluid).

Similarly,

$$G = \frac{1}{2} m V^2 + E'$$

$$= \frac{1}{2} m (\underline{V}(x,t) + \dot{\underline{V}})^2 + E'$$

$$\underline{\underline{Z}} = \int \epsilon \underline{V} f d^3 \underline{V}$$

$$= \int G \underline{V} f d^3 \underline{V}$$

$$\underline{\underline{Z}} = \int d^3 \underline{V} (\underline{V}(x,t) + \dot{\underline{V}}) (G' + \frac{1}{2} m (\underline{V}(x,t) + \dot{\underline{V}})^2) f$$

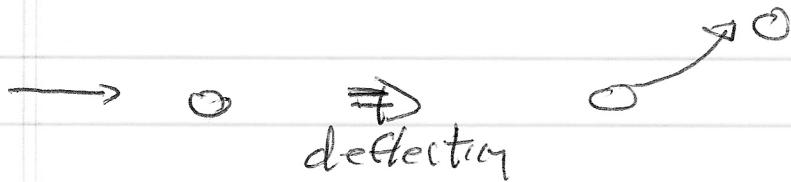
$$= \underline{V}(x,t) \left(\frac{1}{2} m \dot{V}^2 + \underbrace{P + n \bar{\epsilon}}_{\hookrightarrow W \rightarrow \text{internal energy}} \right)$$

Z

$$\frac{\partial}{\partial t} (n \bar{\epsilon}) + \nabla \cdot \left[\underline{V}(x,t) \left(\frac{1}{2} m \dot{V}^2 + P + n \bar{\epsilon}' \right) \right] = 0$$

\downarrow
can simplify
as far momentum.

→ angular momentum not conserved by
 $C(f)$)



→ most truncations stop at 3rd moment

Now, for transport:

→ what if f_0 is gradient
 \Rightarrow driver of f

i.e. Chapman-Enskog expansion

$$\Rightarrow \delta f = -\frac{1}{r} \underline{v} \cdot \underline{D} f_0$$

→ then $f_0 + \delta f$ set fluxes, so

$\text{Flux} = \underset{f_0}{\text{ideal part}} + \underset{\delta f}{\text{collisional part}}$

δf

\Rightarrow transport

Key Point: How relate fluxes to
 gradients in thermodynamic
 quantities?